Roman Piffl / CEMeNt / EDMOND

- * EDMOND contents data about **3 060 250** meteors (31. 12. 2014)
- Combination of that data produce 210 887 orbits of meteors (after reduction by means of qualitative criteria)
- Source data from MetRec and UFO Tools contain not only astrometric information
- In that data we can find also informations about meteor brightness
- ✤ But...
- Data on brightness are not used (except in the case of absolute magnitude)
- ✤ So...
- There are 3 million meteor lightcurves hidden in our database!

✤ How to unhide that LCs?

- 1. Because MetRec and UFO Tools have a different format of data storage, we need two conversion utilities to obtain data on brightness.
- 2. We need to choose optimal format of data storage for further processing.
- 3. We need to create some utilities for combination of multistation meteor lightcurves.
- 4. We need to create tools for combination LCs with data about atmospheric trajectory.
- 5. We need to create a graphical interface for new database to be welcoming to the user.

Format of MetRec data

- MetRec data are in two files
 - 1. *.ref file
 - a) it contains data about site of observation
 - b) data about refences stars
 - 2. *.inf file
 - a) It contains data about *.ref file
 - b) data about captured object
 - i. Time
 - ii. Magnitude
 - iii. RA and DEC

Format of UFO data

- ✤ UFO data are in one XML file
 - 1. It contains detailed information about station, camera and captured object
 - 2. Captured object is determined frame by frame with information about
 - a) Frame number
 - b) Sum of luminosity
 - c) Magnitude
 - d) Azimuth and elevation
 - e) RA and DEC

SiteCode 14270													
Longitude 11.191670													
Latitude 46.144199													
Altitude 680													
OperationMode unguided													
ReferenceDate 2012 4 8													
ReferenceTime 19 30 0													
NoiseLevel 6													
MaxStarDiameter 5.9													
MinStarDiameter 1.0													
VideoBrightness 240													
VideoContrast 255													
CenterOfPlate 192 144													
OrderOfPlateConstants 3													
NumOfRefStars 134													
RefStar1 1.8 1.1 363 11.0620 61.7510 128.75 24.50													
RefStar2 1.8 0.0 259 12.9000 55.9600 46.50 72.00													
RefStar3 3.0 1.6 149 10.3720 41.4990 212.00 123.75													
Refstar4 2 3 0.0 215 11 0310 56 3820 140 25 54 50													

MetRec

*.ref file

AppearanceDate 01.01.2013 AppearanceTime 17:13:53 ReferenceStars 20120408.ref FrameCount 20

#	time brig	ght x	У	alpha	delta	c_x	c_y	c_alpha	c_delta	use
01	53.447	- 0.246	0.735	3.3154	54.752					no
02	53.487	- 0.247	0.717	3.3228	53.935					no
03	53.527	- 0.247	0.700	3.3306	53.119					no
04	53.567 2.	6 0.248	0.682	3.3384	52.304			3.3364	52.328	yes
05	53.607 1.	8 0.249	0.665	3.3462	51.490			3.3455	51.453	yes
0 6	53.647 2.	2 0.249	0.645	3.3535	50.567			3.3544	50.576	yes
07	53.687 2.	0 0.250	0.627	3.3599	49.729			3.3629	49.699	yes
Ø 8	53.727 1.	9 0.250	0.607	3.3726	48.813			3.3711	48.820	yes
0 9	53.767 2.	0 0.251	0.588	3.3774	47.909			3.3791	47.940	yes
10	53.807 2.	1 0.252	0.569	3.3872	47.042			3.3868	47.059	yes
11	53.847 1.	8 0.252	0.550	3.3970	46.159			3.3943	46.176	yes
12	53.887 2.	0 0.253	0.532	3.4019	45.321			3.4015	45.292	yes
13	53.927 2.	1 0.254	0.513	3.4082	44.434			3.4085	44.407	yes
14	53.967 2.	2 0.256	0.493	3.4126	43.493			3.4153	43.531	yes
		250								

 \wedge

MetRec

*.inf file

*.XML file

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UFO Analyser

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dct="-12.037234" memo="">																		
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cdegmax="0.033879" io="3" raP="352.863861" dcP="-9.014144"
av1="4.958315" x1="317.848053" y1="117.213600" x2="339.450806"
y2="84.878815" az1="6.649399" ev1="26.623760" az2="9.405540"
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dc1="67.902039" ra2="281.802765" dc2="63.951656" ram="283.466858"
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dist2="197.310410" gd2="180.078674" len="15.383389" GV="19.229235"

MetRec brightness determination

 In the first detection step, a mean background image is subtracted from the digitized frame. The mean image is derived from the previous frames.
 Negative pixel values in the difference image are set to zero, so that only pixels with increased brightness are further examined.

UFO brightness determination

The magnitude of one field is computed from light sum of the pixels of the object in one field. Light sum is the sum of the increased light of pixels of the object.

Increased light = brightness of the pixel at the instance – brightness of the pixel before the appearance of the object.

Background is mean of 10 frames before.

✤ After conversion, we can create database with following format:

- 1. Station (latitude, longitude, altitude)
- 2. Camera (FOV, sensor, resolution etc.)
- 3. Informations about each frame
 - a) Date and time
 - b) Lsum (only from UFO data)
 - c) Magnitude
 - d) Azimuth and elevation
 - e) RA and DEC
- 4. Image of meteor (only from UFO data?)

Problems with obtained data

- 1. No informations about reduction we need to use that data "as it is"
- 2. No reduction for following effects:
 - a) distance
 - b) speed
 - c) vignetting and field distortion
 - d) meteor parts (shock waves, body, foot)
 - e) in frames of video data are often lacking owing to the impossibility of detecting peak
 - f) background noise
 - g) saturation (above 255 level)
 - h) spectral sensitivity
 - i) extinction, refraction, the effect of height above the horizon at the signal level (increased background etc.)
 - j) light pollution

Benefits of that database

- 1. Big data for statistics analyses
- 2. In combination with EDMOND database:
 - a) Determination of typical LCs for more than 300 regular showers
 - b) Determination of typical LCs for outbursts
 - c) Comparison of LCs year by year and filament by filament for big showers
 - d) Investigation of preheating phase for high altitude meteors
 - e) Investigation of fragmentation and physical parameters for different types of meteoroid bodies
 - f) Etc.

Current status

- 1. Data collection for additional data mining
- 2. Creating procedures for populating the database (for UFO it is done)
- 3. Creating GUI for database usability
- 4. Testing of charting normalized curves (with data of outburst of DRA and SPE)
- 5. Searching for a suitable name of LCs database (working name is **METaLICA** METeor Light Curve dAtabase)

✤ Timetable

- 1. End of 2015 conversion procedures will be done
- 2. Spring 2016 database in testing mode
- 3. IMC 2016 presentation of database

Normalized LC of Draconid



Normalized LC of Draconid with flash



Any questions?



Thank you for your attention!

Roman Piffl / CEMeNt / EDMOND / METALICA